

# Long-Baseline Neutrino Oscillations

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IV International Workshop on the Interconnection  
Between Particle Physics and Cosmology

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# Introduction

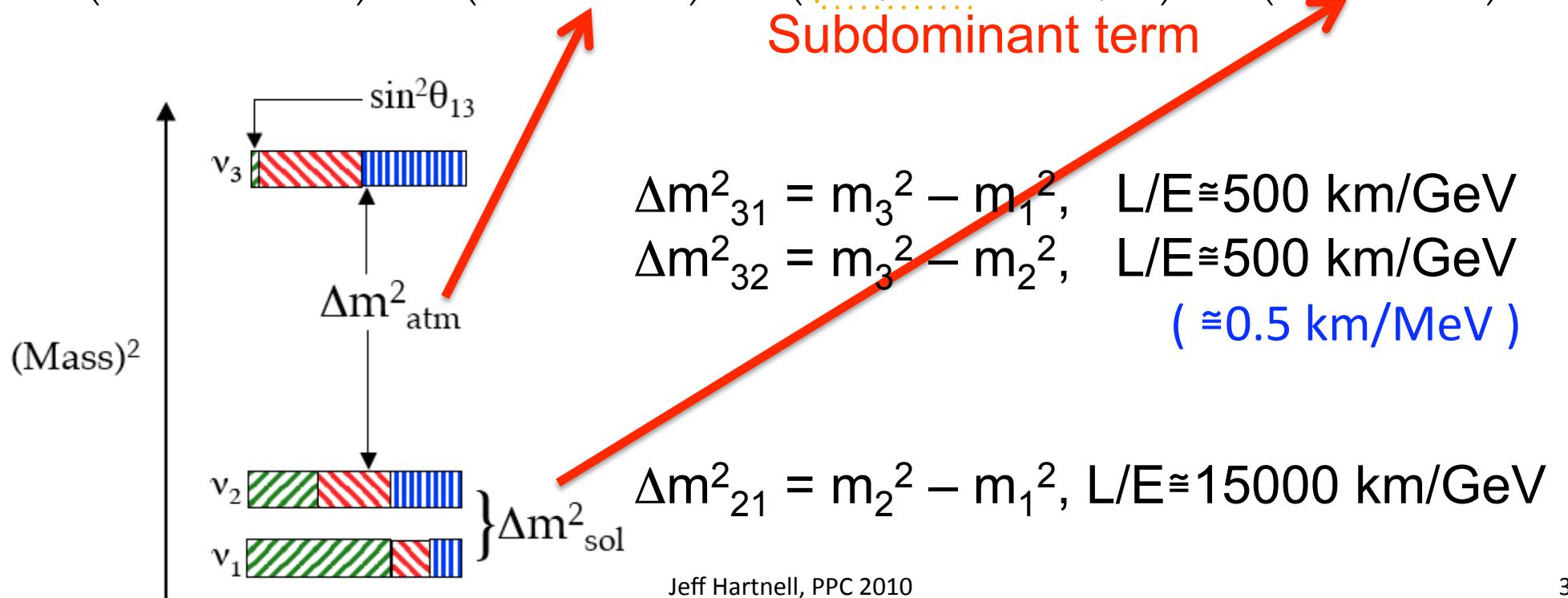
- Theory overview
- Where we stand experimentally
- Results:
  - Solar sector: Solar + KamLAND
  - Atmospheric sector: MINOS + SK
  - Subdominant sector: Chooz + MINOS
- Future:
  - Reactor expts (**Double Chooz**, Daya Bay and RENO)
  - Accelerator expts (NOvA)
    - T2K and LBNE (see following talks)

# Theory Overview

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

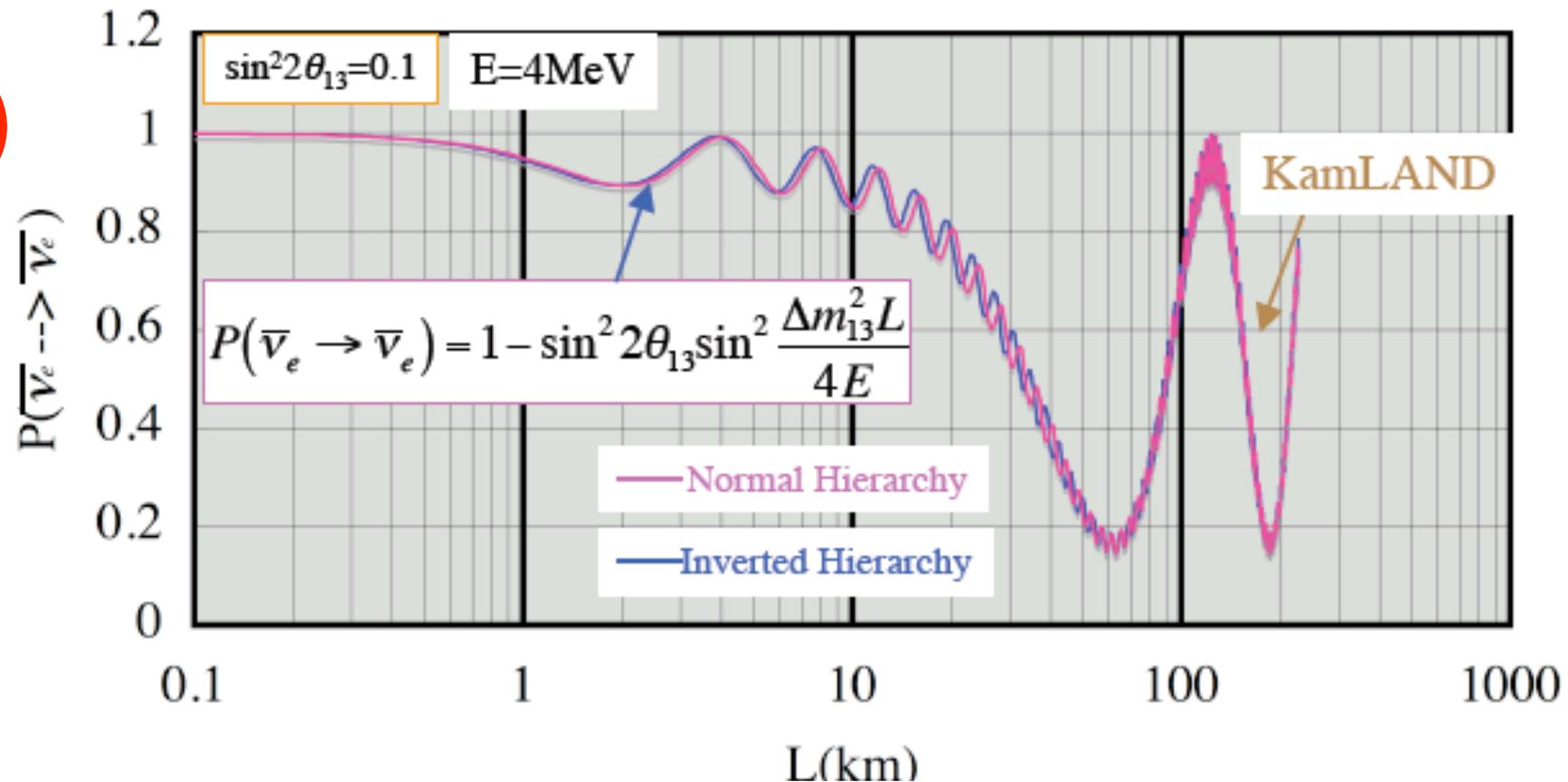
$$\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{-i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

**Subdominant term**



# Two Examples of Oscillation Probability

1



$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2(2\theta_{13}) \sin^2(\theta_{23}) \sin^2 \left( 1.27 \Delta m_{31}^2 \frac{L}{E} \right) +$$

2

$$\sin^2(2\theta_{12}) \cos^2(\theta_{23}) \sin^2 \left( 1.27 \Delta m_{21}^2 \frac{L}{E} \right) +$$

$$\sin(2\theta_{13}) \sin(2\theta_{23}) \sin(2\theta_{12}) \sin \left( 1.27 \Delta m_{31}^2 \frac{L}{E} \right) \sin \left( 1.27 \Delta m_{21}^2 \frac{L}{E} \right) \cos \left( 1.27 \Delta m_{32}^2 \frac{L}{E} \pm \delta_{CP} \right)$$

# Where we stand experimentally

bf  $\pm 1\sigma$

acc.@ $3\sigma$

$\Delta m^2_{21}$   $(7.7 \pm 0.2) \cdot 10^{-5} \text{ eV}^2$

(8%)

KamLAND

$\sin^2 \theta_{12}$   $0.31 \pm 0.015$

(18%)

SNO

$|\Delta m^2_{32}|$   $(2.4 \pm 0.1) \cdot 10^{-3} \text{ eV}^2$

(20%)

MINOS

$\sin^2 \theta_{23}$   $0.46 \pm 0.045$

(38%)

SK atm

$\sin^2 \theta_{13}$   $< 0.044$  @  $3\sigma$

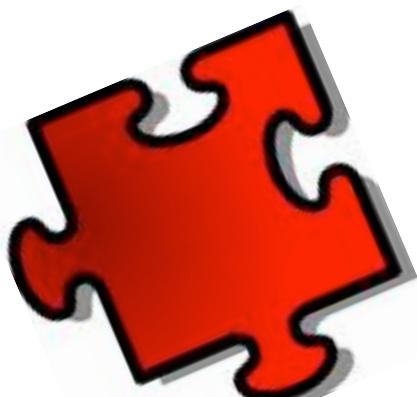
CHOOZ

$\delta_{CP}$ ,  $\theta_{23}$  quadrant, mass hierarchy (?)

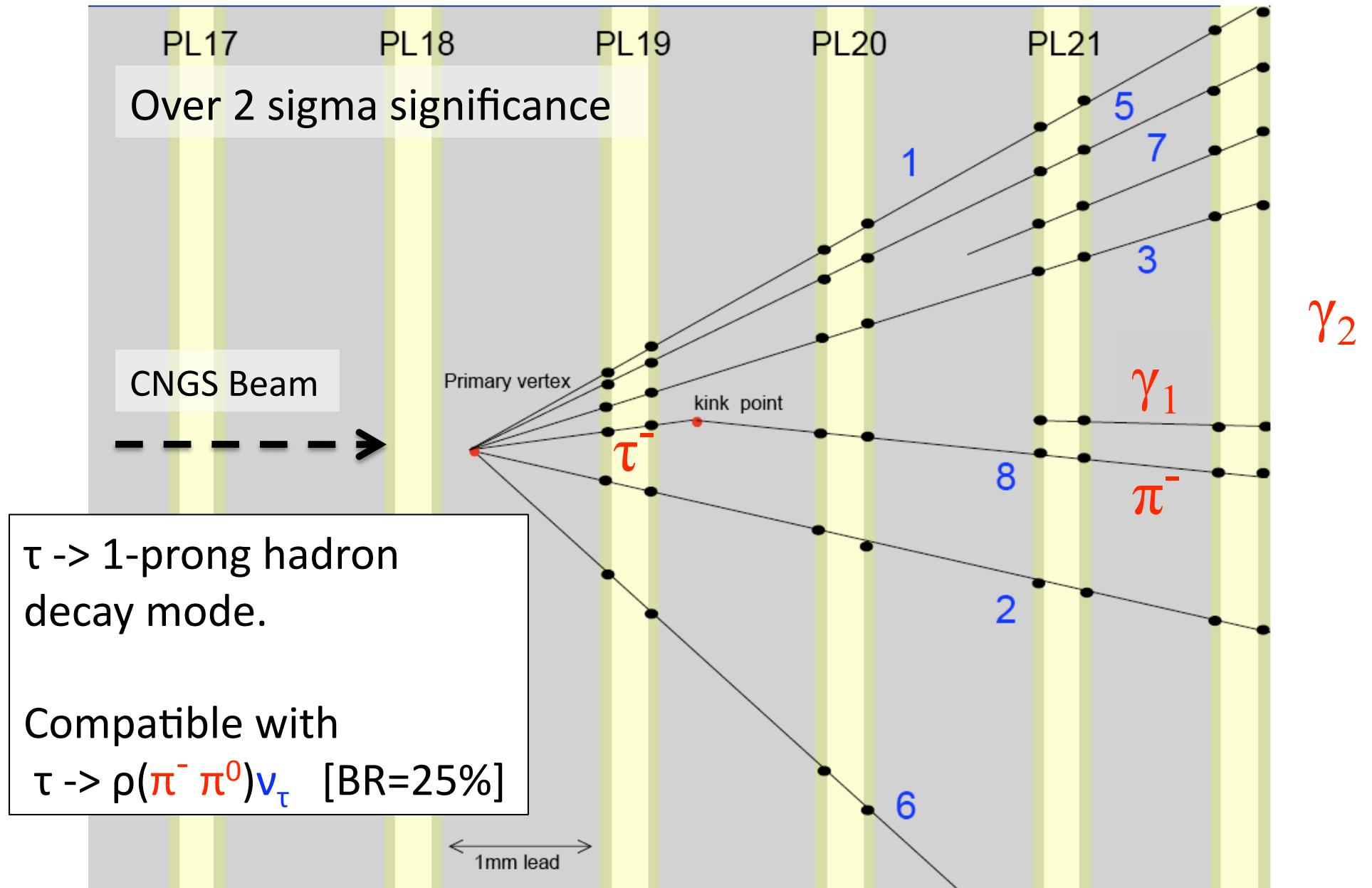
# First direct observation of a neutrino flavour appearing due to oscillations

(neglecting LSND/MiniBooNE)

## Verifying the oscillation model...



# Opera's 1<sup>st</sup> $\nu_\tau$ Candidate Event



# Results breakdown

Solar sector: Solar+KamLAND

Atmospheric sector: MINOS+SK

Subdominant sector: Chooz + MINOS

# Solar Sector

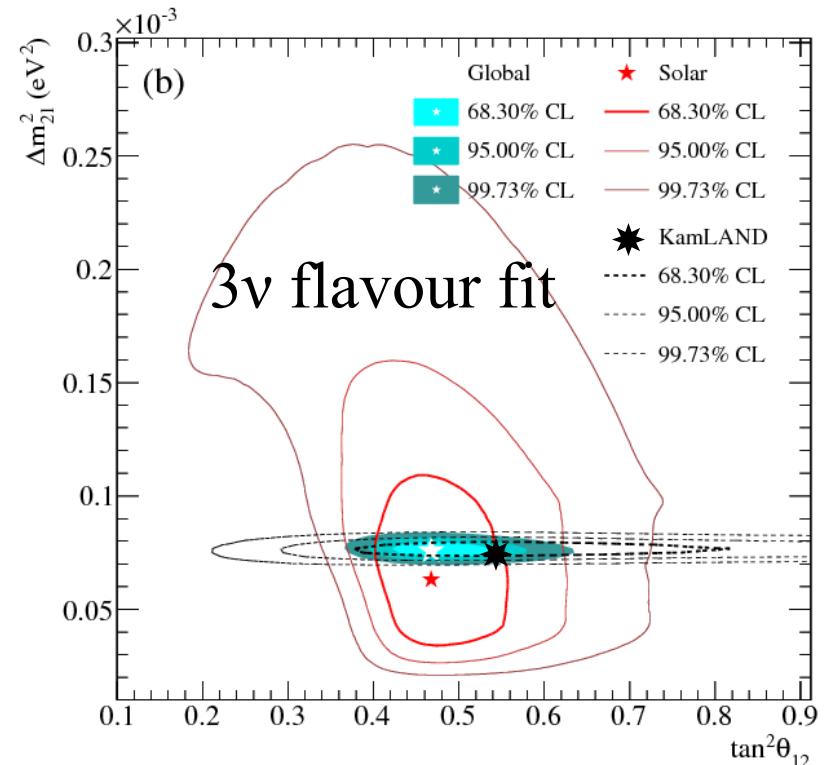
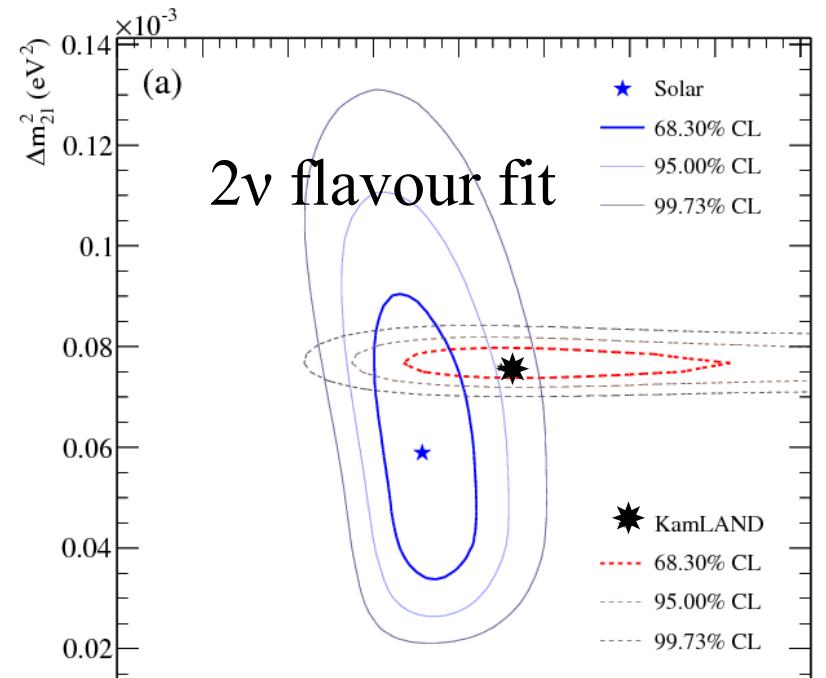
# Solar + KamLAND

- SNO Low Energy Threshold Analysis (LETA)
  - some improvements to come
- SK, Chlorine, Gallium, Borexino expts included
- KamLAND (data: 2002-2007)
  - wont double again until 2013(?)
  - $\Delta m^2$  largely systematically limited

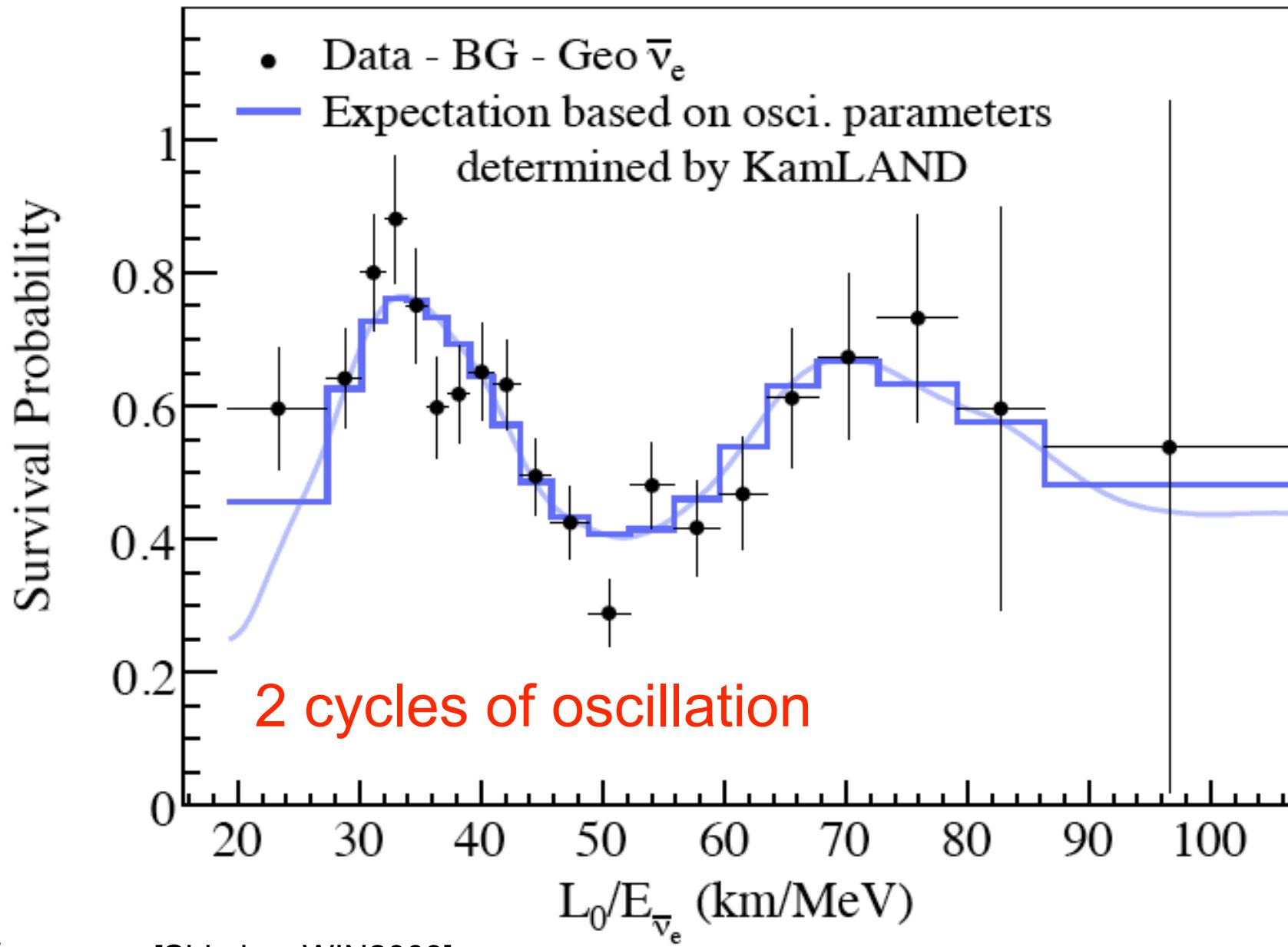
3v flavour best-fit:

$$\sin^2 \theta_{13} = 0.020^{+0.021}_{-0.016}$$

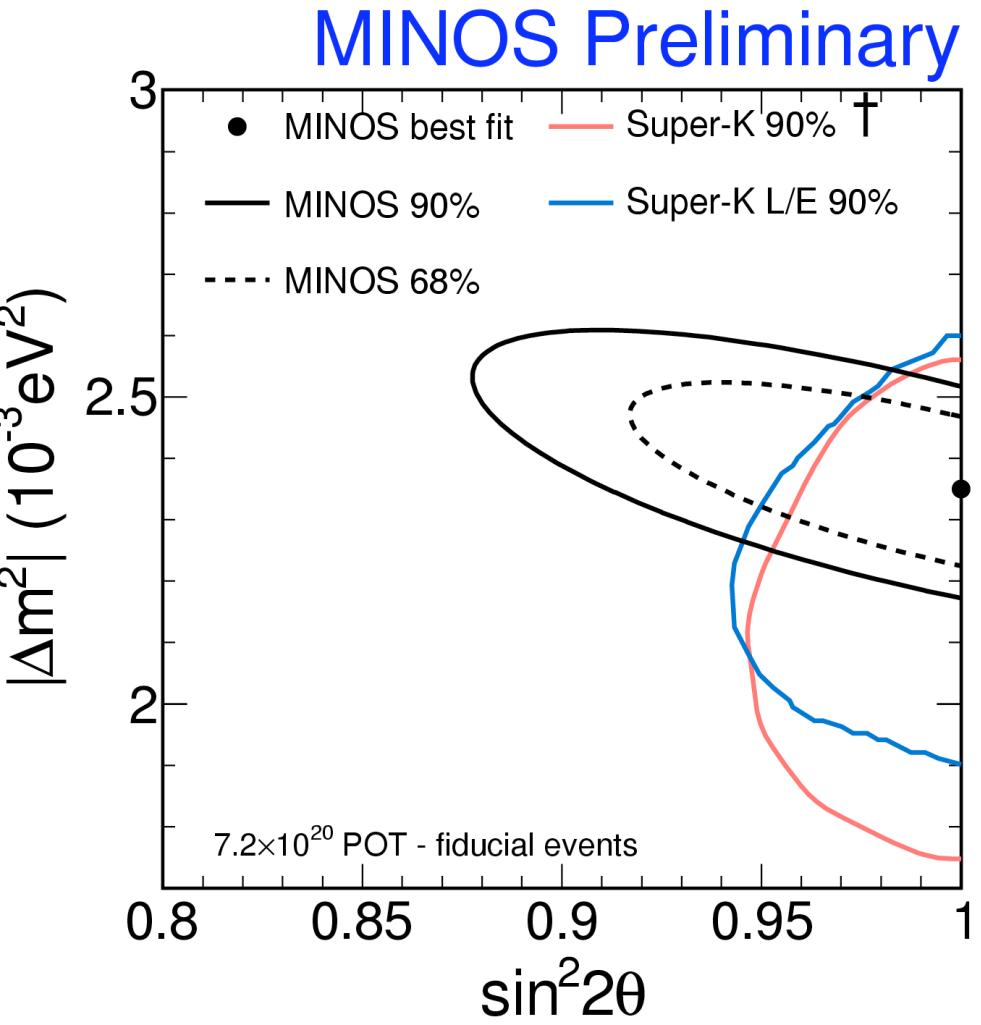
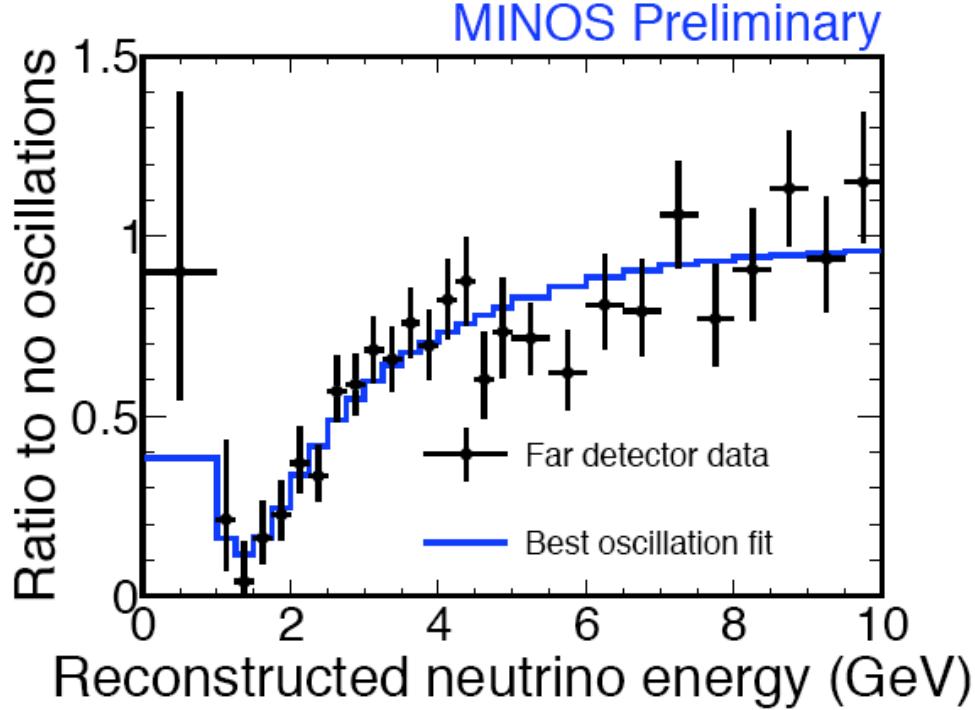
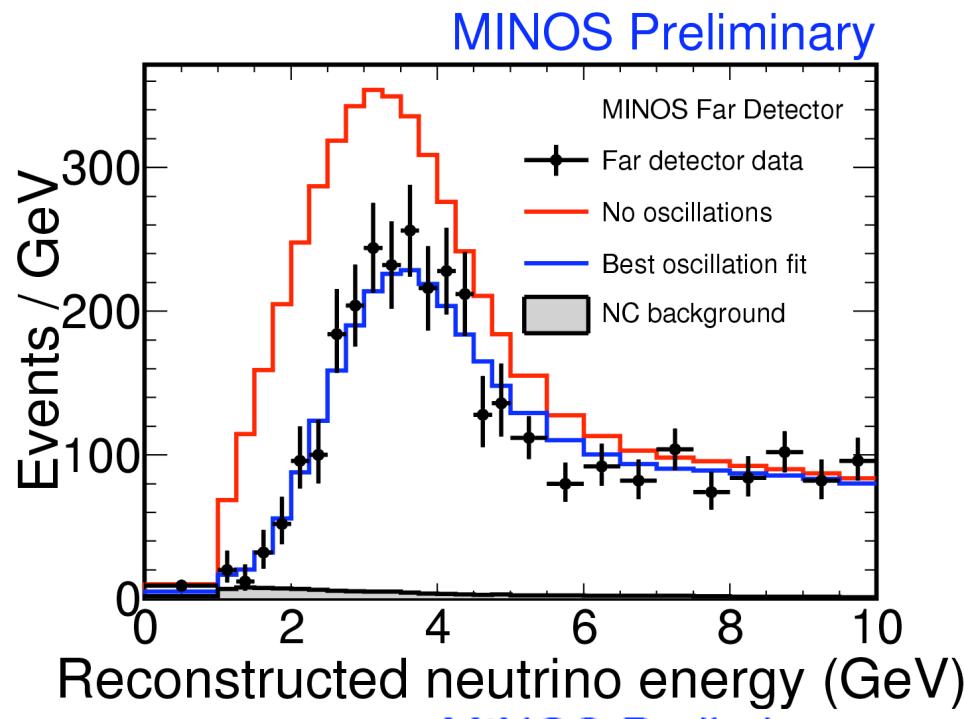
$$\sin^2 \theta_{13} < 0.057 \text{ (95% C.L.)}$$



# KamLAND's Beautiful Data



# Atmospheric Sector



$$|\Delta m^2| = 2.35^{+0.11}_{-0.08} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\theta) > 0.91 \text{ (90\% C.L.)}$$

†Super-Kamiokande 2-flavour (preliminary)

# Full 3-flavor oscillation results



SK-I+II+III Preliminary

May 2010

Super-K has extended from 2-flavour fit to 3-flavour, includes

- $\delta_{CP}$
- $\theta_{13}$
- matter effect
- fixed solar osc. ( $\Delta m^2_{21}$ ,  $\theta_{12}$ )

SK is sensitive due to:

- mixed flux of  $\bar{\nu}_e \bar{\nu}_\mu \bar{\nu}_e \bar{\nu}_\mu$
- large range of L, E and L/E

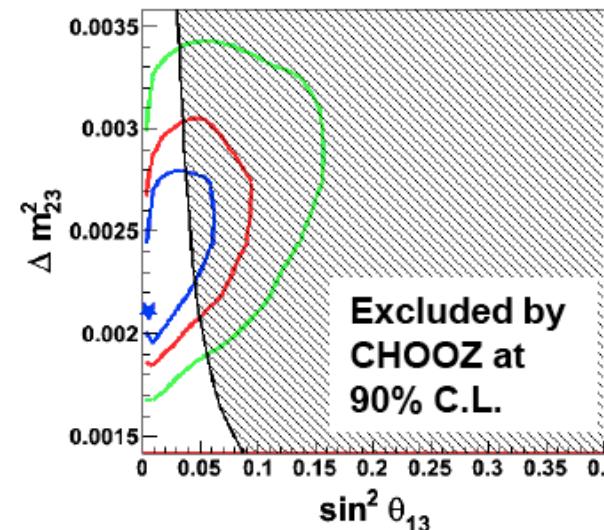
PRD81, 092004, 2010

Considers only matter and solar terms individually (neglects interference term between the two)

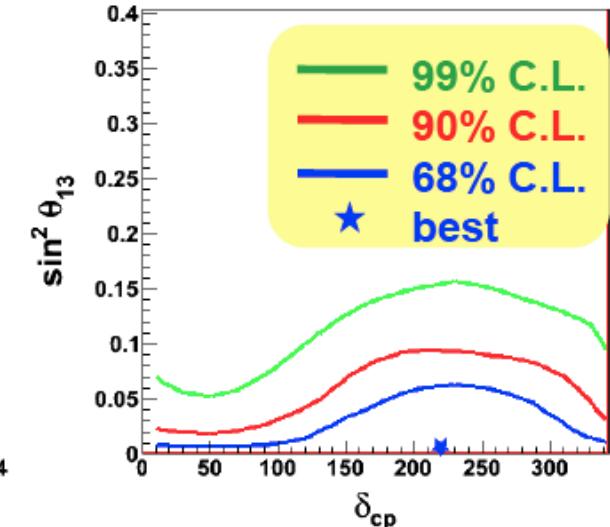
New FULL 3-flavour analysis this summer

- No significant preferences on  $\theta_{13}$ , mass hierarchy or  $\delta_{CP}$

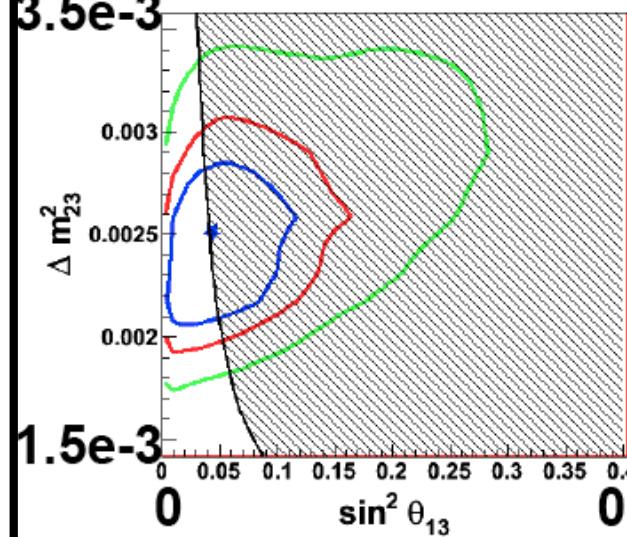
Normal hierarchy



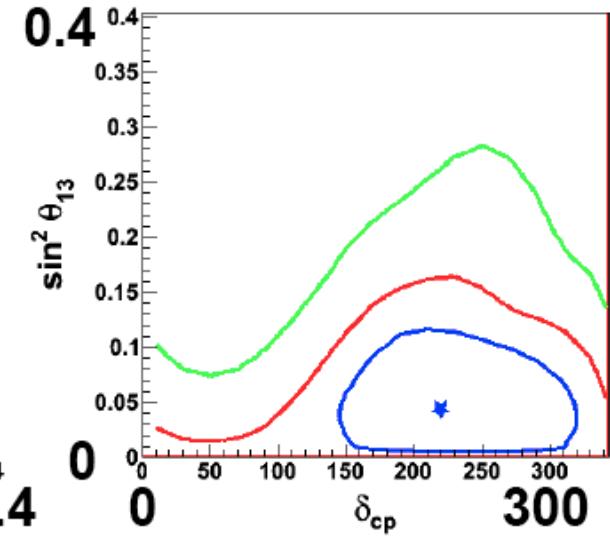
Normal hierarchy



Inverted hierarchy



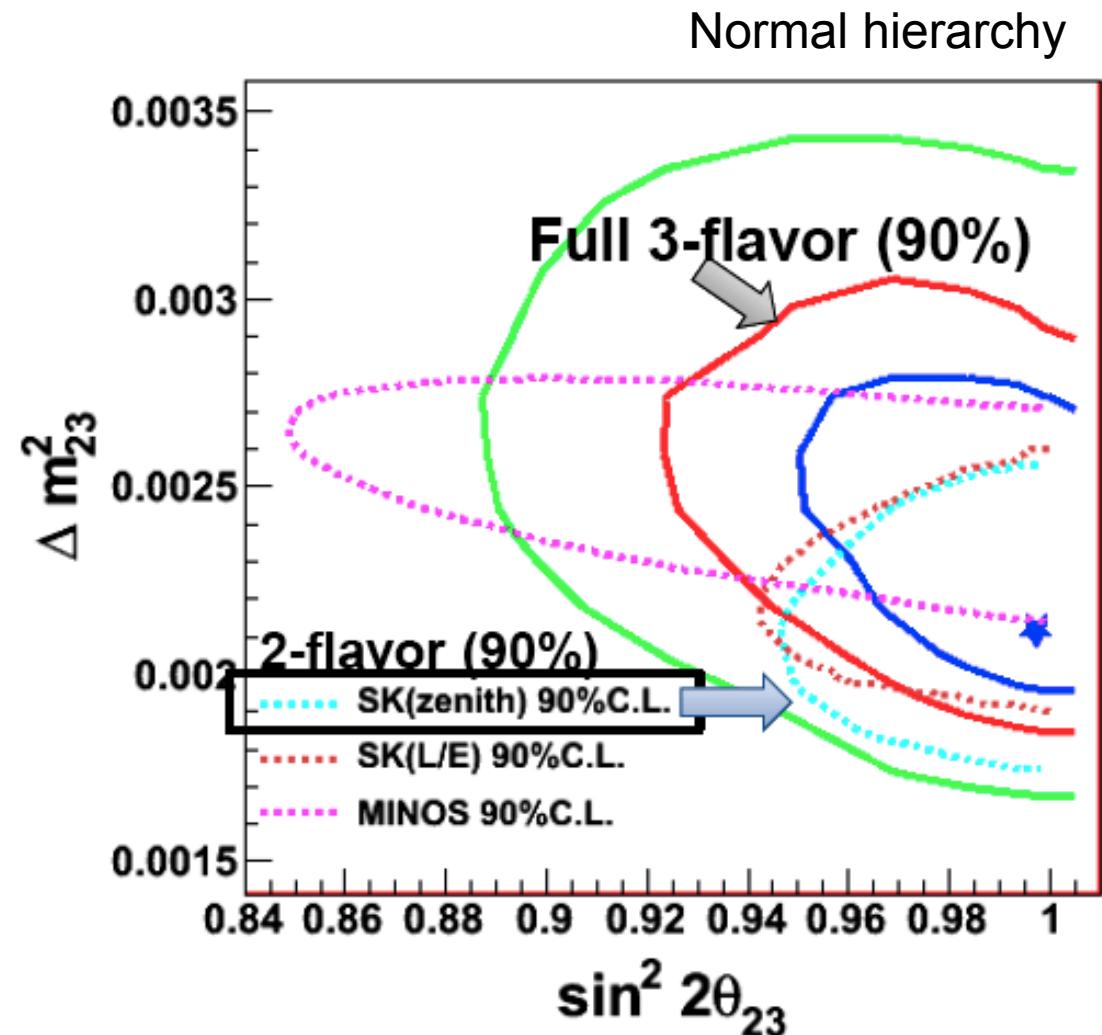
Inverted hierarchy



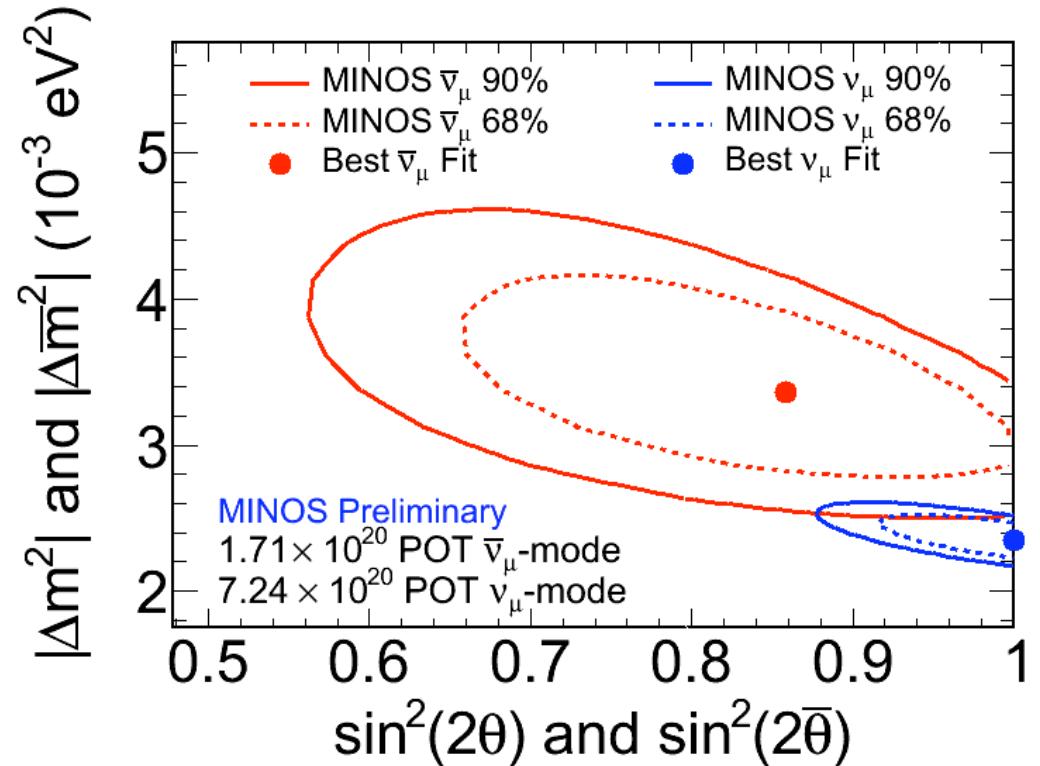
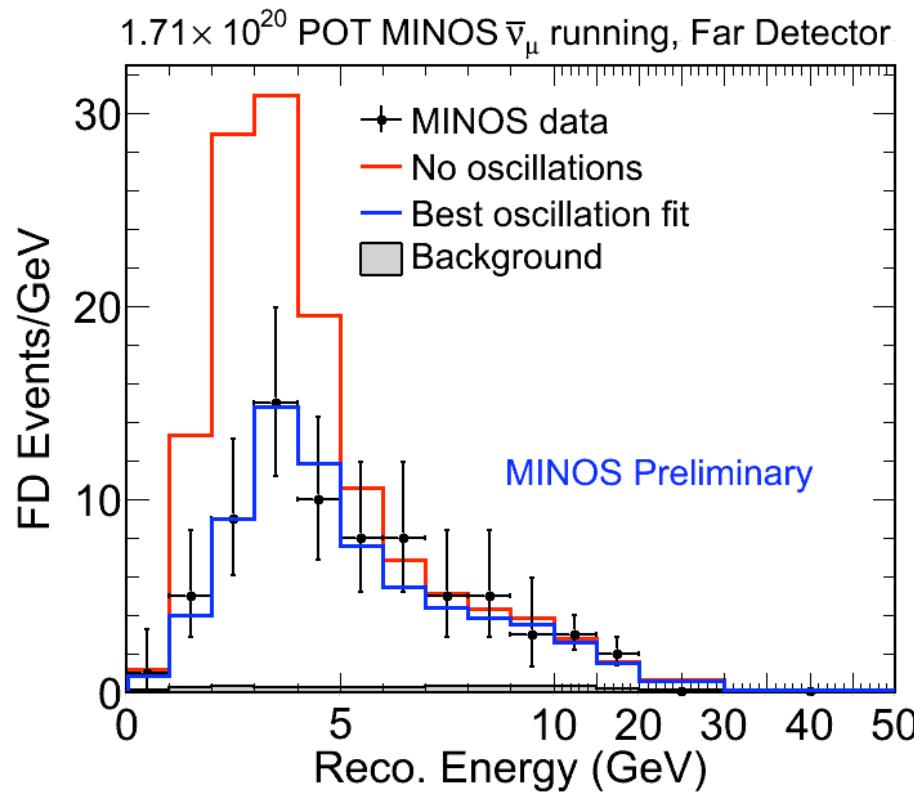
# Comparing 3-flavour with 2-flavour

Including  $\delta_{CP}$ ,  $\theta_{13}$  and matter effect weakens SK constraints on the “2-flavour parameters”

Naively, not expected to make a large difference to MINOS contour (note: old contour shown)



# Comparison of $\nu_\mu$ and $\bar{\nu}_\mu$



6.3  $\sigma$  rejection of  $\bar{\nu}_\mu$   
null-oscillation hypothesis

$$|\Delta m^2| = 3.36_{-0.40}^{+0.45} \times 10^{-3} \text{ eV}^2$$

$$\sin^2(2\bar{\theta}) = 0.86 \pm 0.11$$

# Subdominant Sector

# MINOS

## $\nu_e$ Appearance and CHOOZ

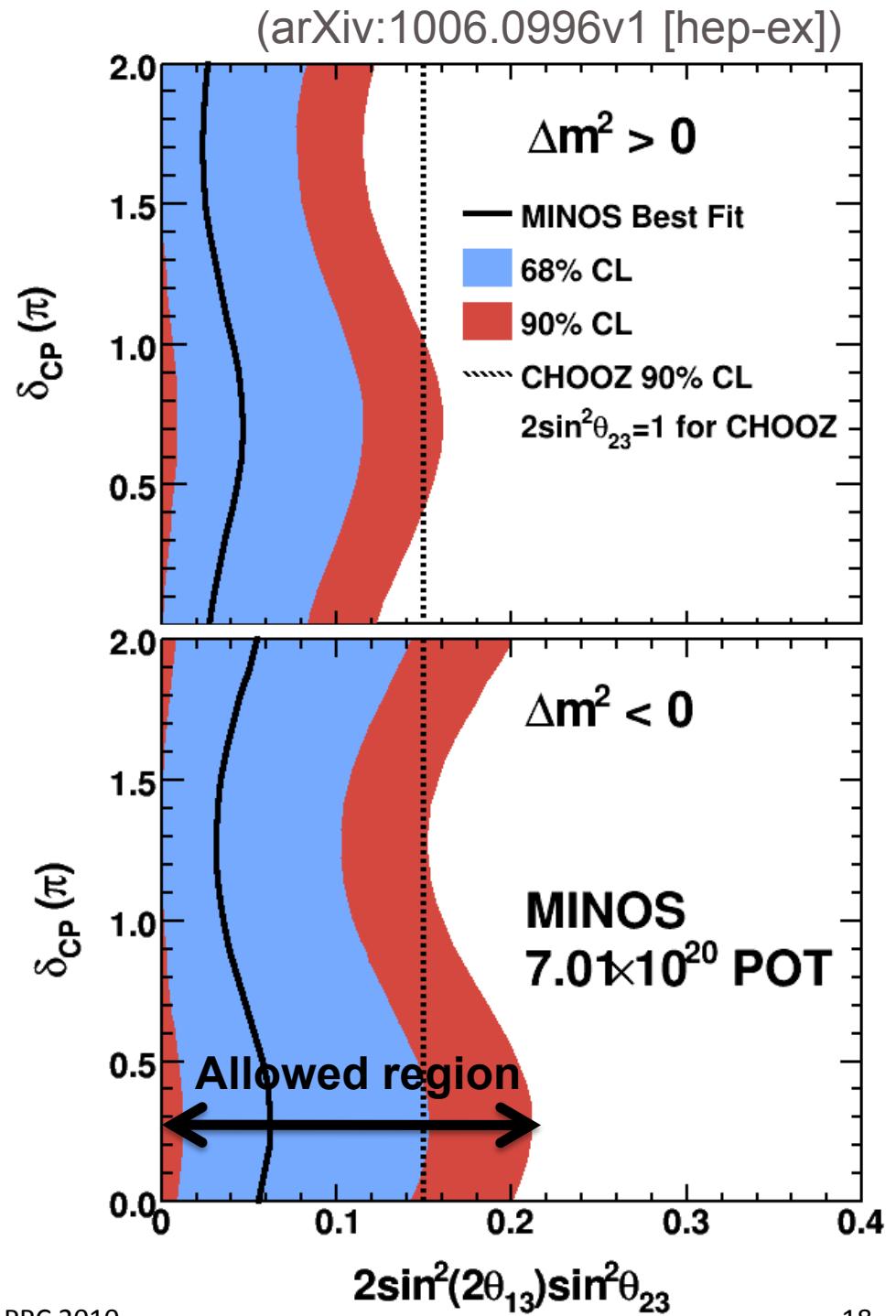
## $\bar{\nu}_e$ Disappearance

Accelerator and reactor experiments are complementary

Accelerator searches for  $\nu_e$  appearance are sensitive to 4 parameters:

$\theta_{13}$ , mass hierarchy,  $\delta_{CP}$  and  $\sin^2\theta_{23}$

Reactor experiments just sensitive to  $\theta_{13}$



# The Future

Hottest topic:

$$\theta_{13}$$

If not too small, provides a potential door to  
leptonic CP violation  
(and determines future experimental program)

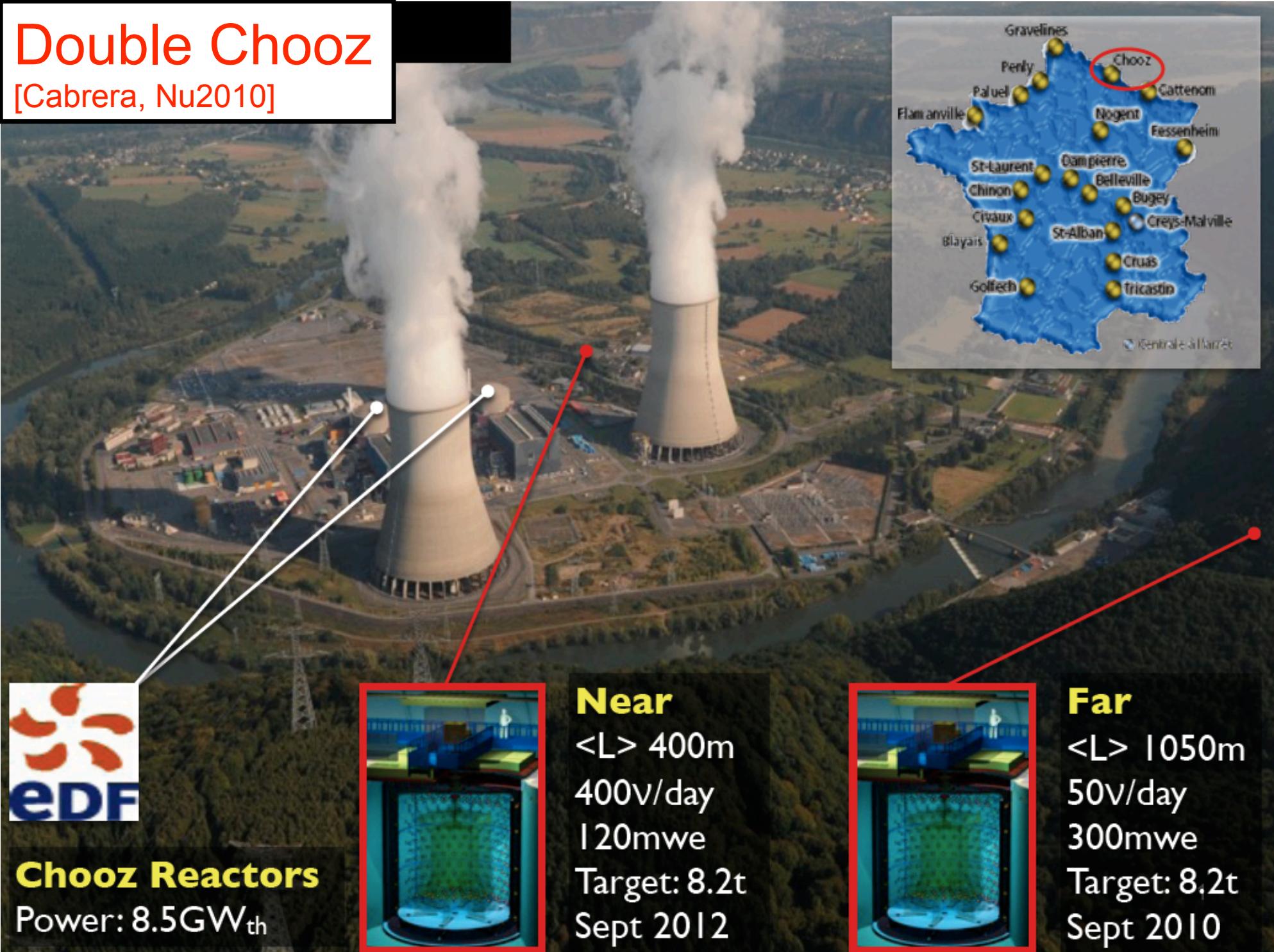
# Dedicated experiments

Accelerator:  
T2K, NOvA, (LBNE)

Reactor:  
Double Chooz, Daya Bay, RENO

# Double Chooz

[Cabrera, Nu2010]



## Energy-Trigger

large energy deposition over  
~70ns coincidence  
(now still without scintillator)

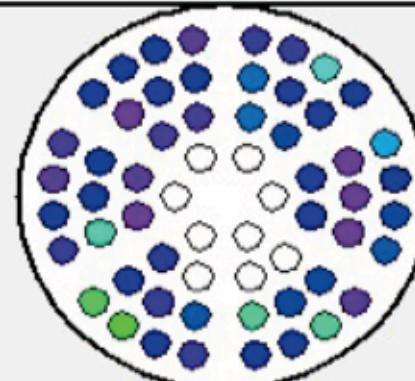
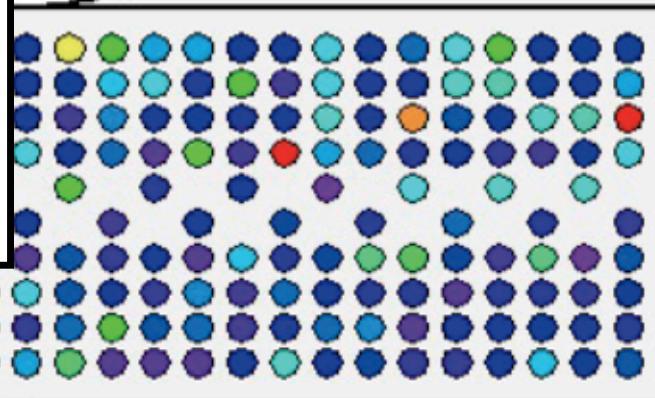
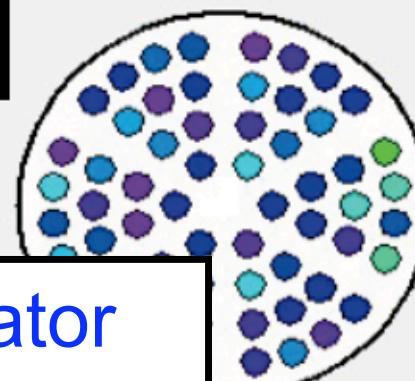
first internal triggers...

Far detector scintillator  
filling to start this month

Neutrinos by September!

Most electronics up &  
running!  
(one FADC card missing)

PRELIMINARY



[Cabrera, Nu2010]



RS-IN2P3 & APC

# A look ahead...

# “Predictions are always hard, especially about the future” (anon)

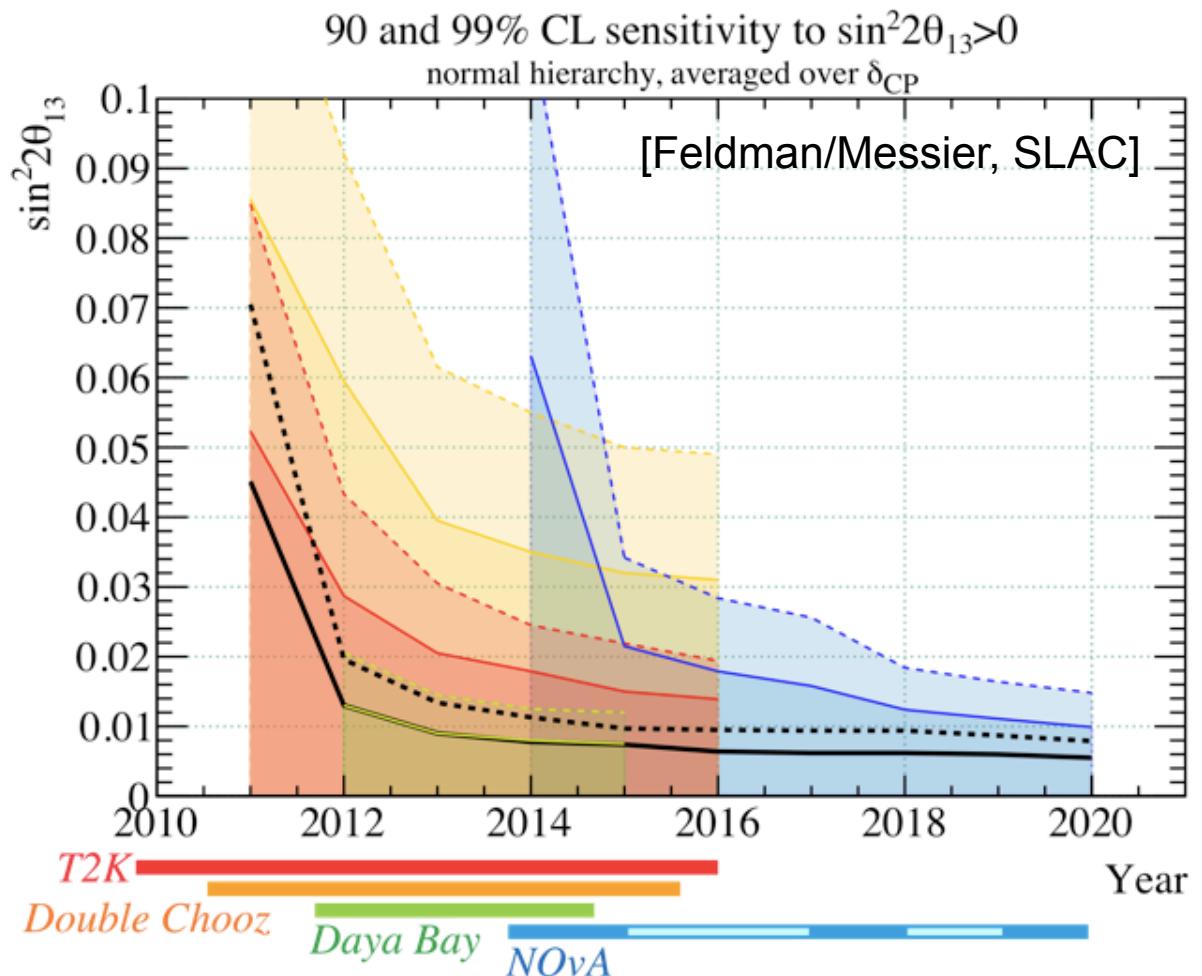
All these experiments could potentially be delayed by a year or more in getting their results out: who knows!

Accelerator searches tend to be optimistic and quote the best or average limit on  $\theta_{13}$

Reactor experiments have a clean measurement of  $\theta_{13}$

Safe to say that 2011-2015 will be an exciting time...

By the end of it we will know if  $\sin^2(2\theta_{13}) > \text{few-}\%$



# Conclusions

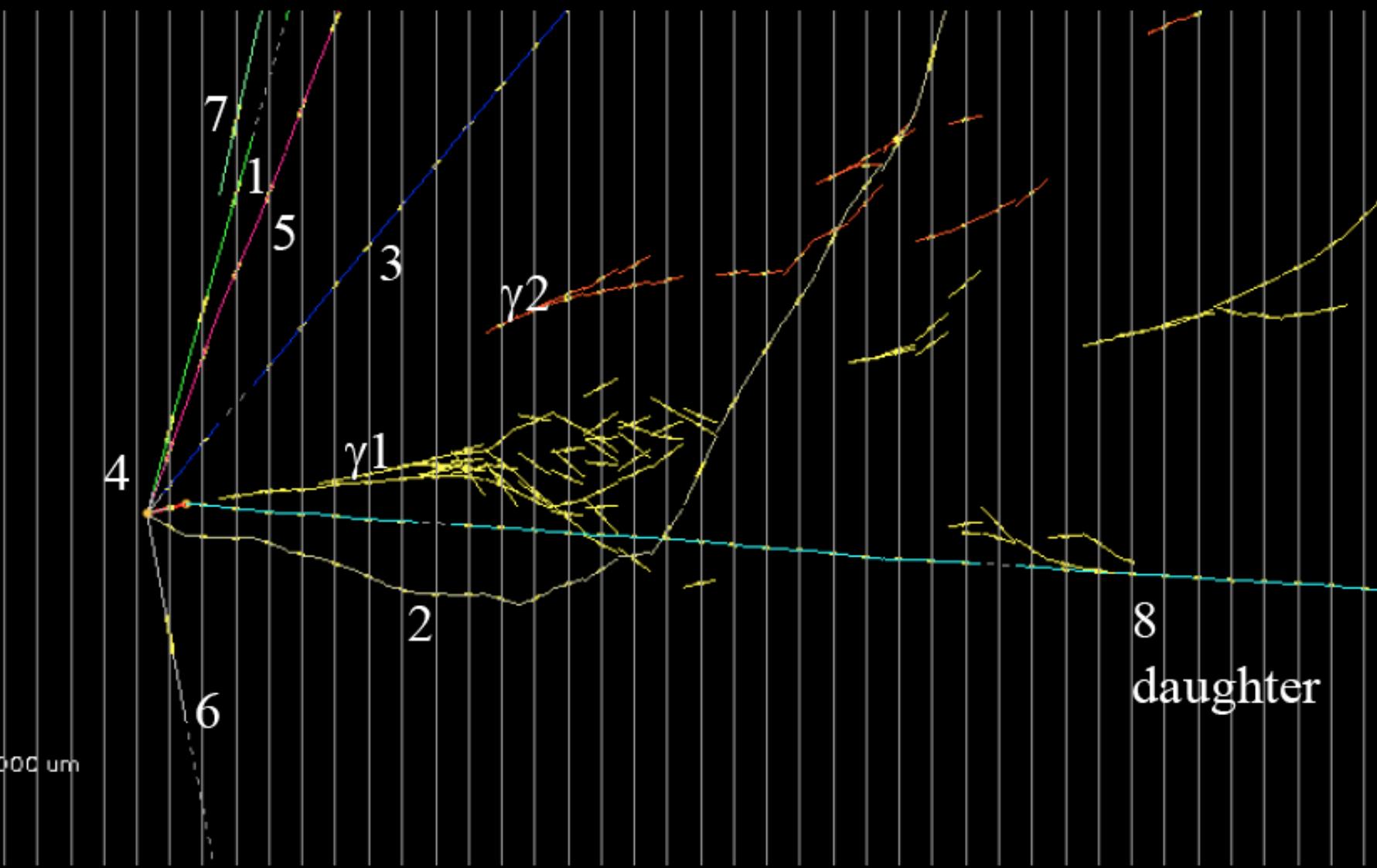
- Diverse neutrino sources + detectors have produced beautiful data
- Precision (~%) measurements of some osc. pars.
  - Reaching the need for full 3-flavour analyses, often can't rely on **accidental bifurcation** of oscillations into **solar and atmospheric sectors**
  - yet several parameters remain undiscovered
- Keep testing assumptions: **new surprises!?**
- Hunt for **subdominant osc.** ( $\theta_{13}$ ) fast ramping up as dedicated experiments come online
  - Next 5 years will provide a results deluge
    - it's going to be very exciting!

Thank  
You

# Backup slides

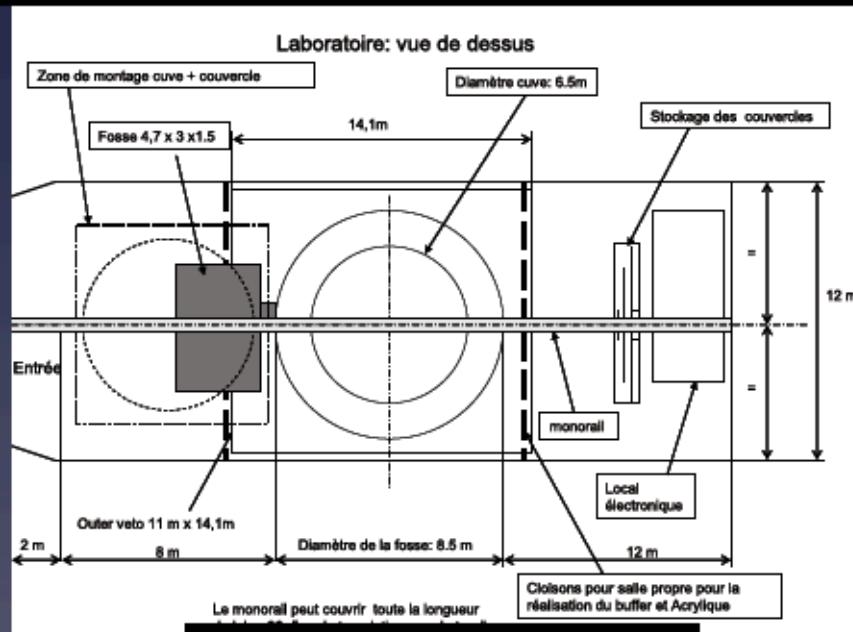
# Event topological features (side view)

Side view



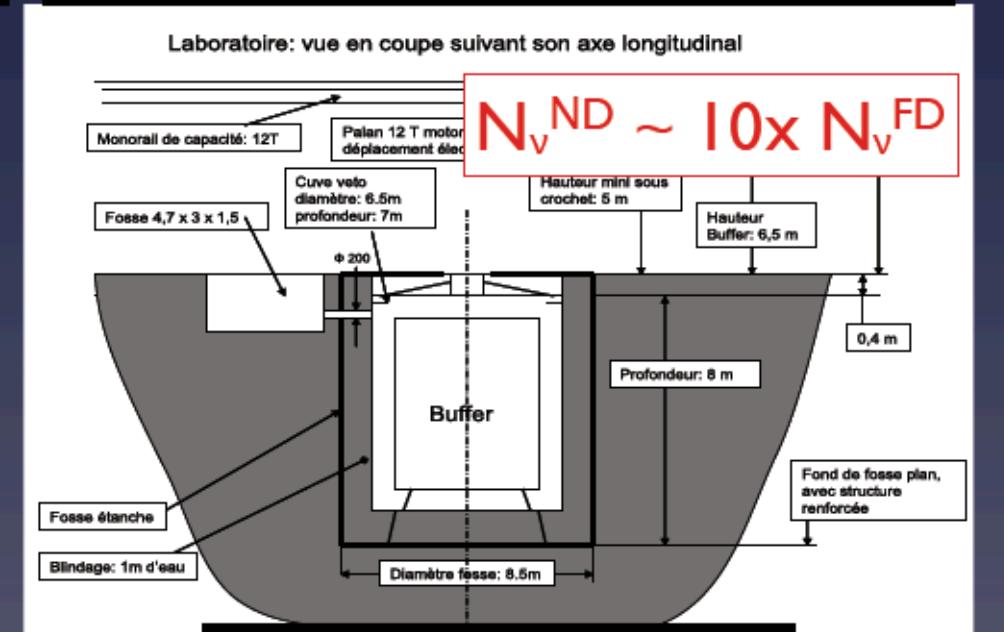
- all authorisations & funding secured...
  - ground studies completed since end of 2009
  - involvement of EDF on the construction
- tendering over summer ⇒ **digging from November 2010**
- lab available fall 2011
- ND construction ⇒ **30 weeks** (many lessons learned from FD)
- goal: **data taking by towards 2012**

## ND Lab designed for building ND



3x zones concept

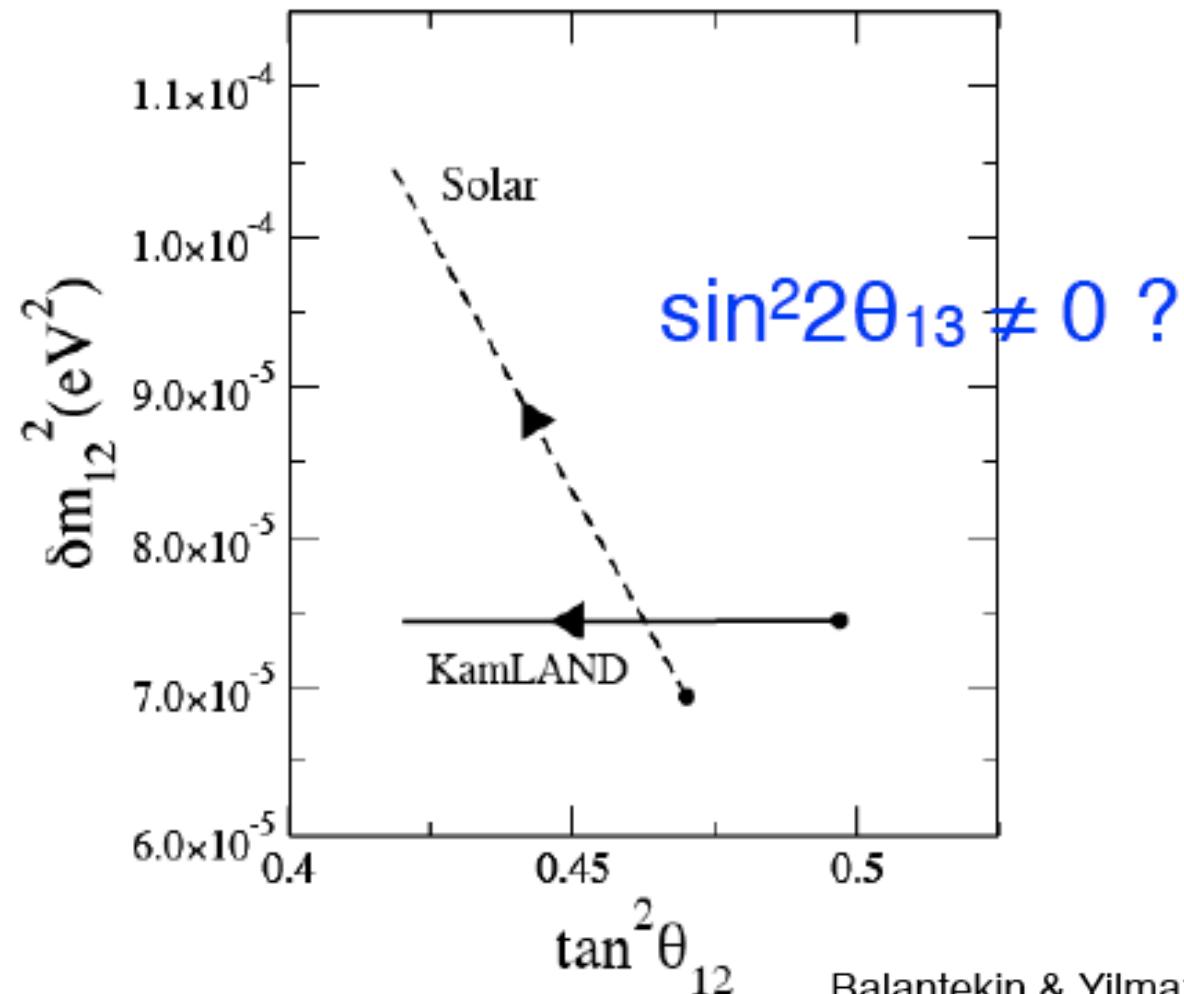
## 1m water shield to radioactivity



$N_v^{\text{ND}} \sim 10 \times N_v^{\text{FD}}$

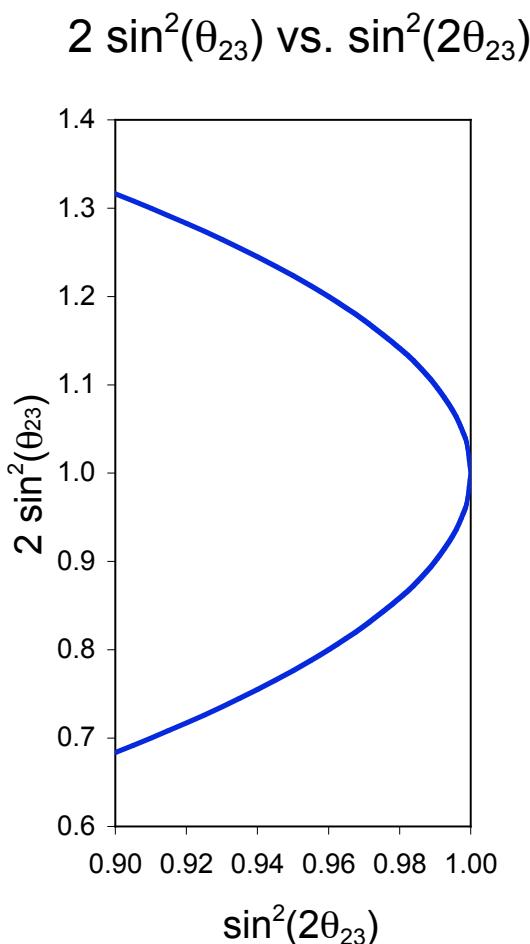
easier filling concept  
(IN2P3 & APC)

# Solar vs. Reactor Tension: effect of theta13



Balantekin & Yilmaz, J. Phys. G **35**, 075007  
(2008) (arXiv:0804.3345 [hep-ph]).

# Effect of not knowing the theta<sub>23</sub> quadrant



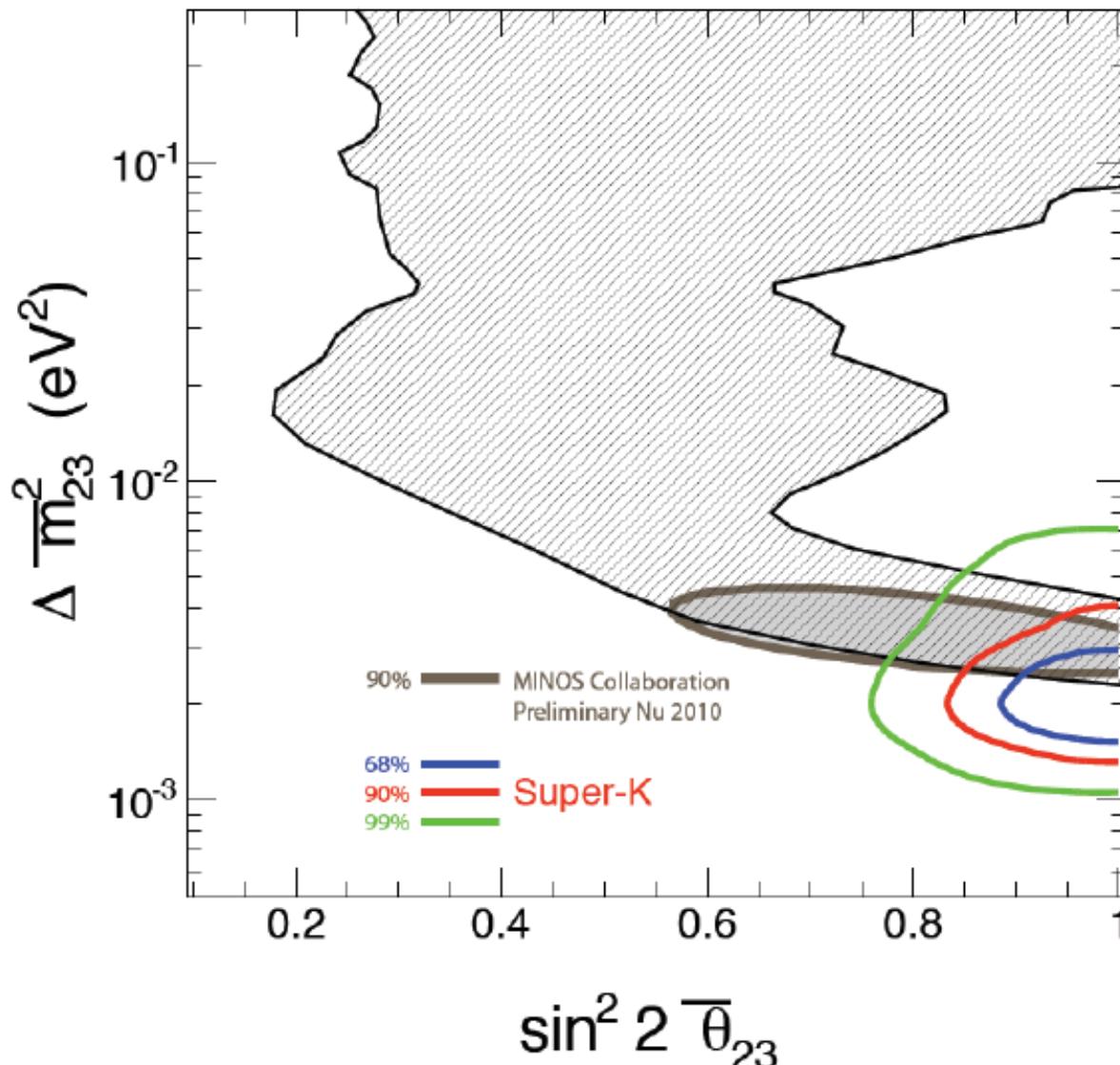
# Search for CPT violation in atm. $\nu$



- Under the CPT theorem,  $P(\nu \rightarrow \nu)$  and  $P(\bar{\nu} \rightarrow \bar{\nu})$  should be same.
- Test  $\nu$  oscillation or  $\bar{\nu}$  oscillation separately.

Jun 2009

SK-I+II+III  
Preliminary



Neutrino:

$$\Delta m_{23}^2 = 2.2 \times 10^{-3} \text{ eV}^2$$
$$\sin^2 2\theta_{23} = 1.0$$

Anti-neutrino:

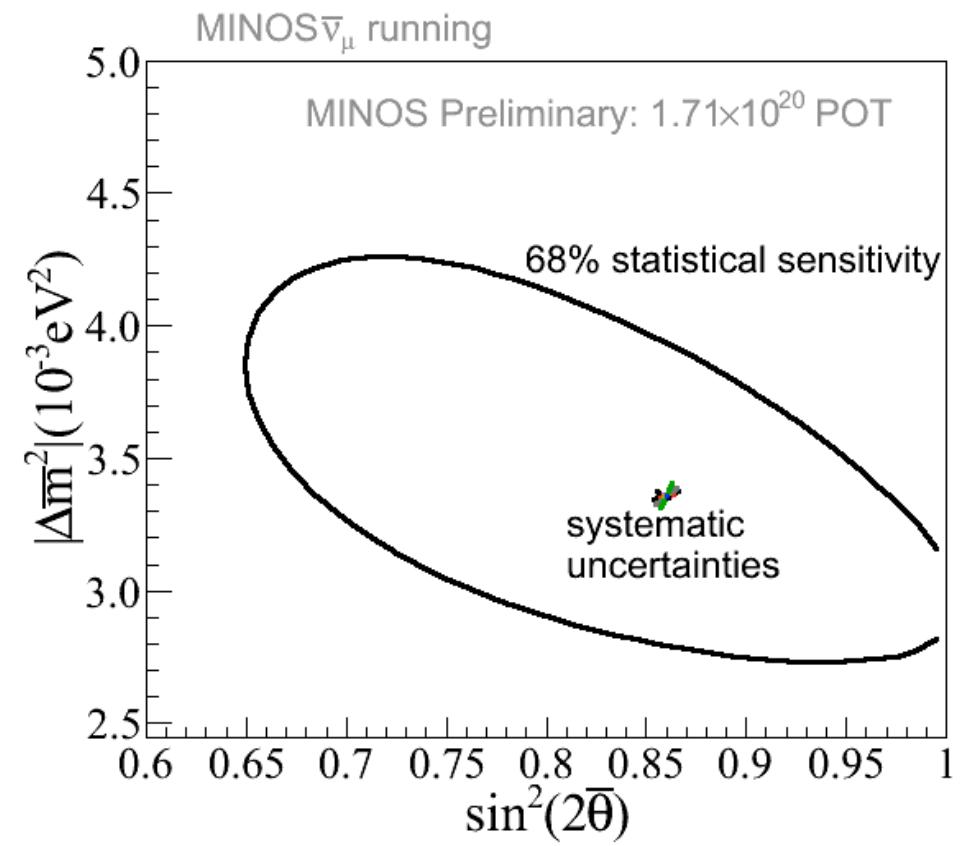
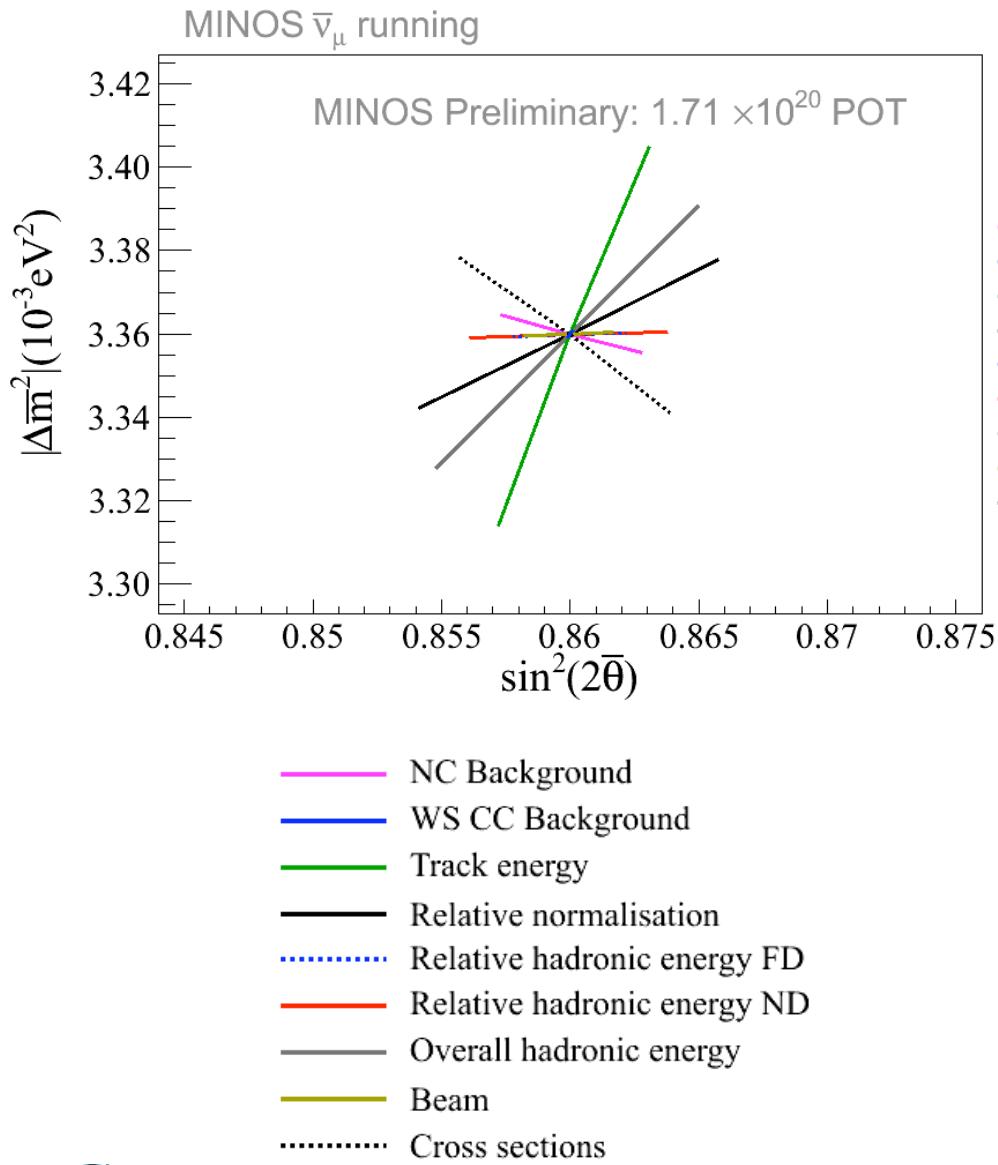
$$\overline{\Delta m}_{23}^2 = 2.0 \times 10^{-3} \text{ eV}^2$$
$$\sin^2 2\overline{\theta}_{23} = 1.0$$

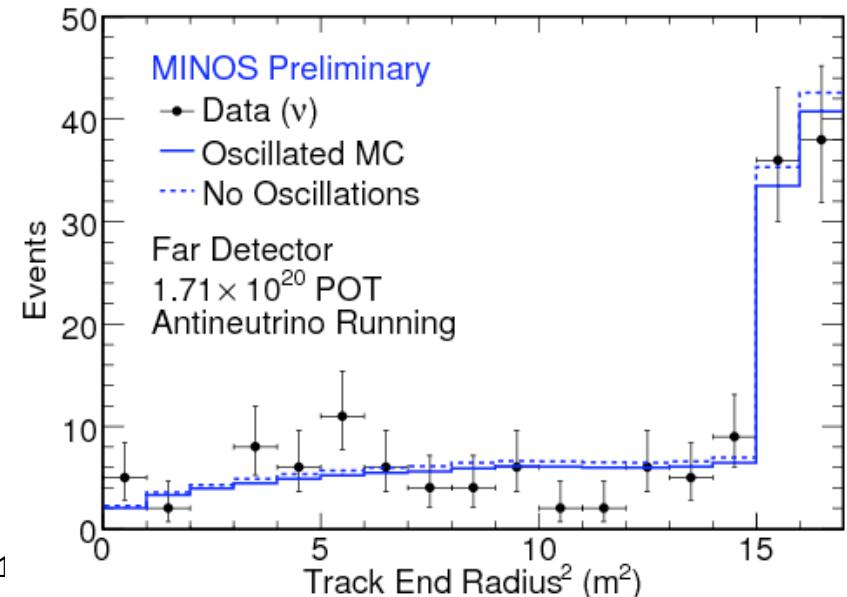
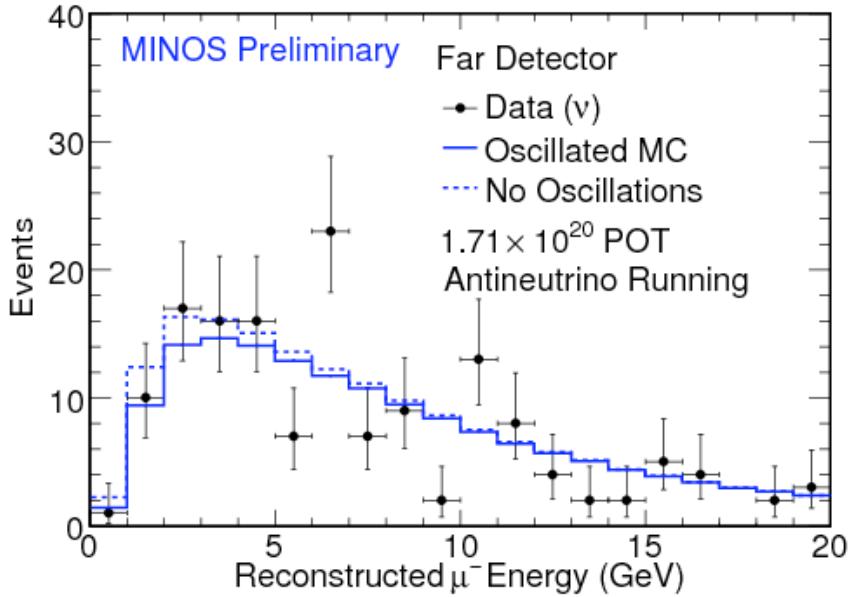
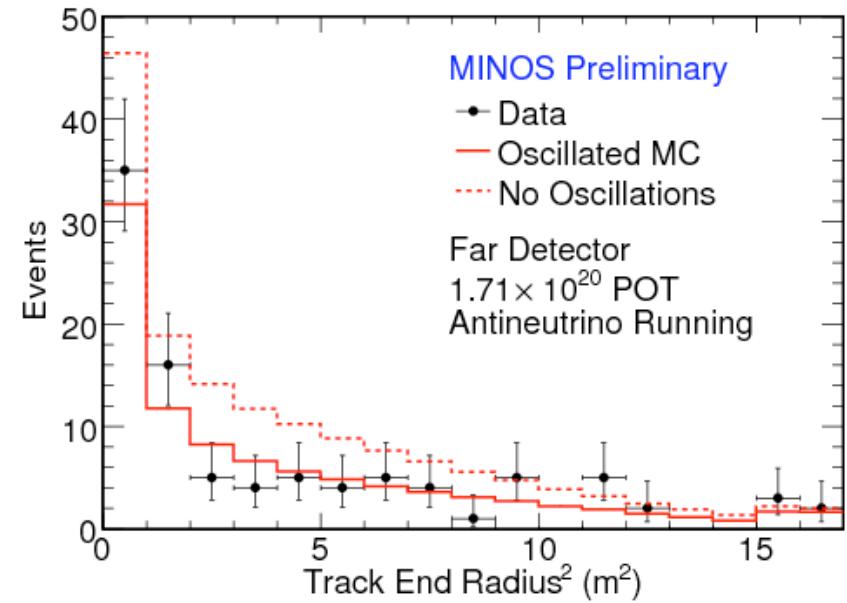
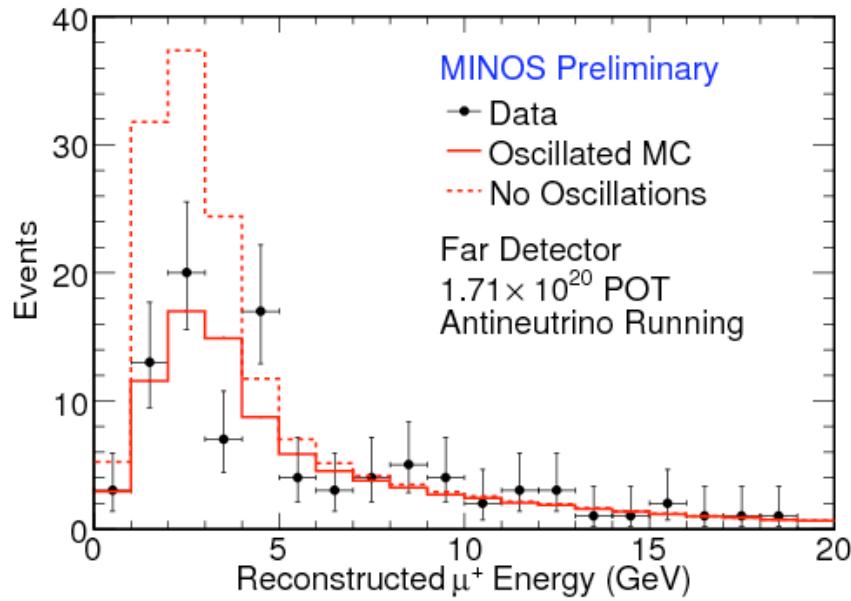
No evidence for CPT violating oscillations is found



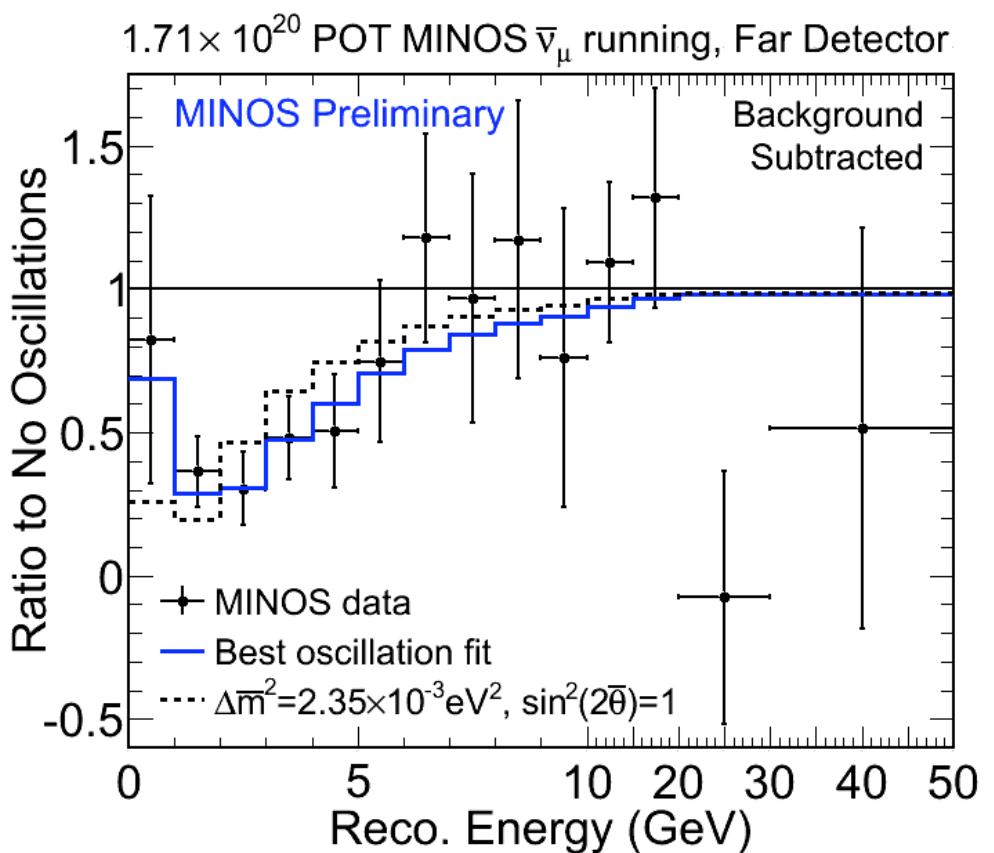
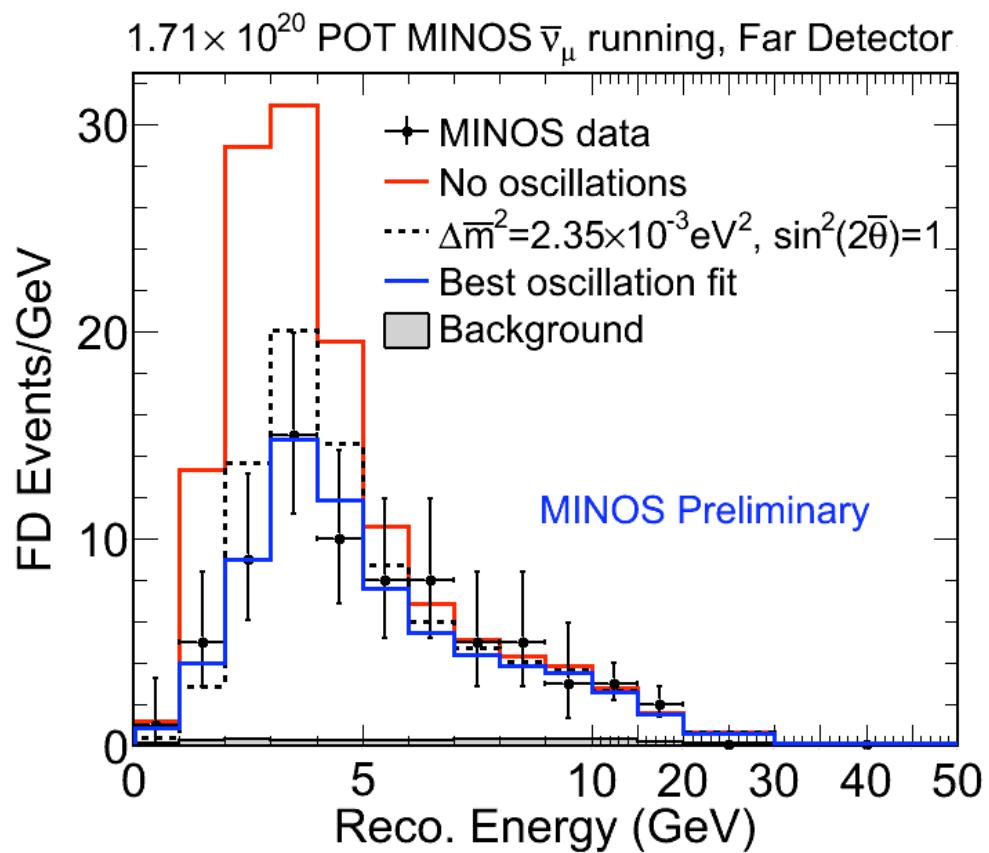
Poster-79 by Roger Wendell

[Takeuchi, Nu2010] 15

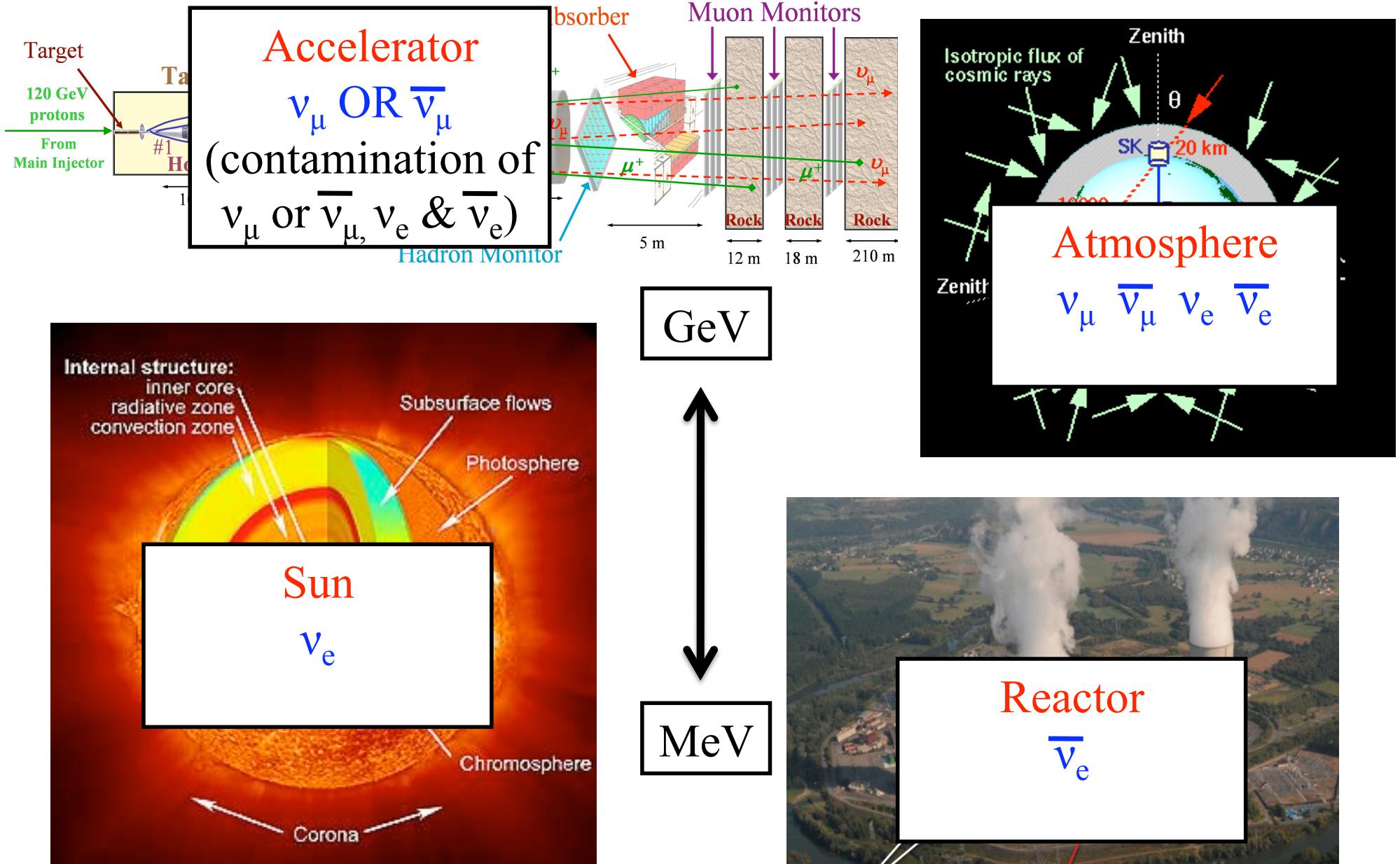




Jeff Hartnell, PPC 201



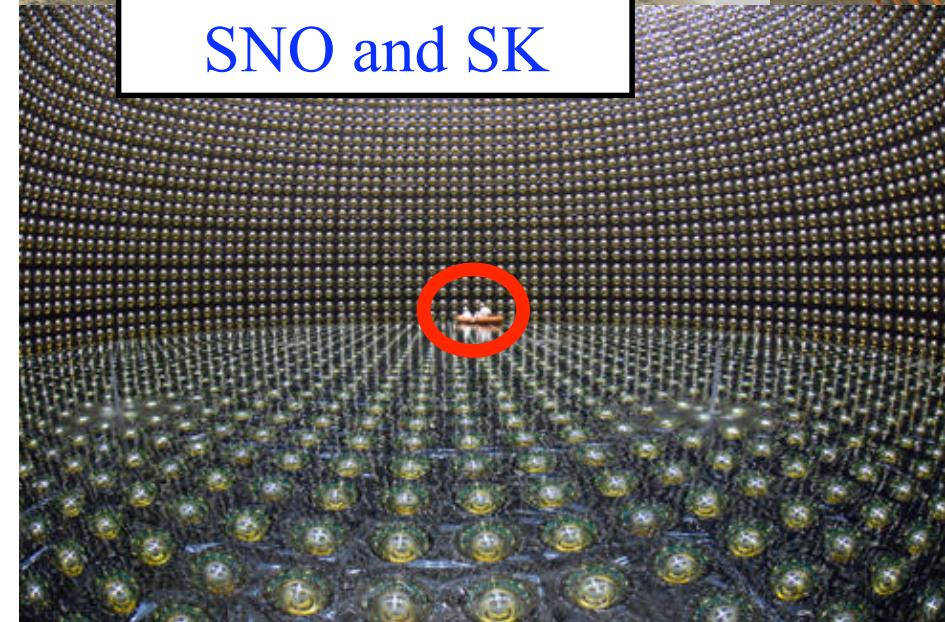
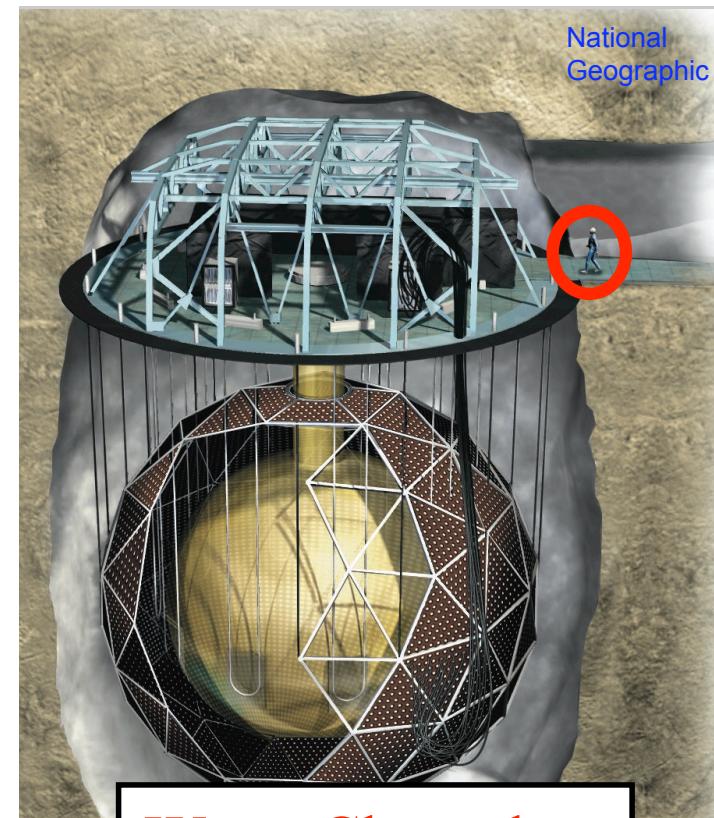
# Neutrino Sources



# Neutrino Detectors

- Massive (Fiducial volume: 10->22000 tons)
  - Shielded/overburden
- Particle Identification:
  - Distinguish  $\nu$ -flavours, e.g.:
    - $\nu_e$  from  $\nu_\mu$
    - $\nu_\mu$  from  $\bar{\nu}_\mu$
    - $\nu_\tau$  from  $\nu_e$  and  $\nu_\mu$
    - Exploit being below kinematic thresholds (e.g. reactor experiments, few MeV)
  - Distinguish CC from NC
  - Distinguish charged lepton from hadrons
- Neutrino/visible energy, direction/position

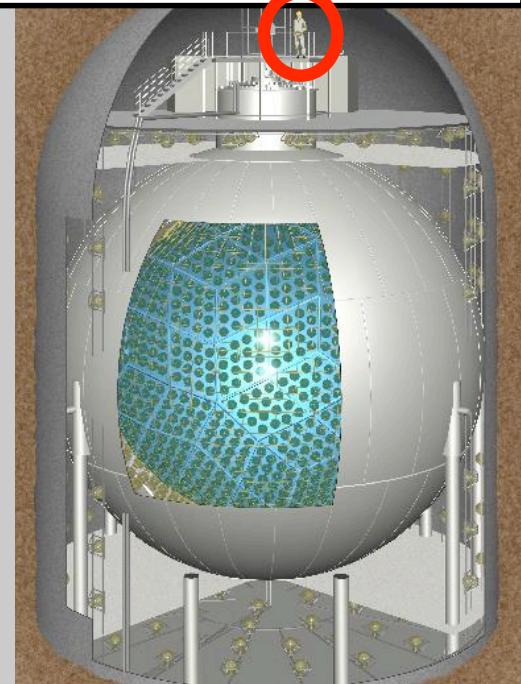
## Iron/Plastic Scintillator MINOS



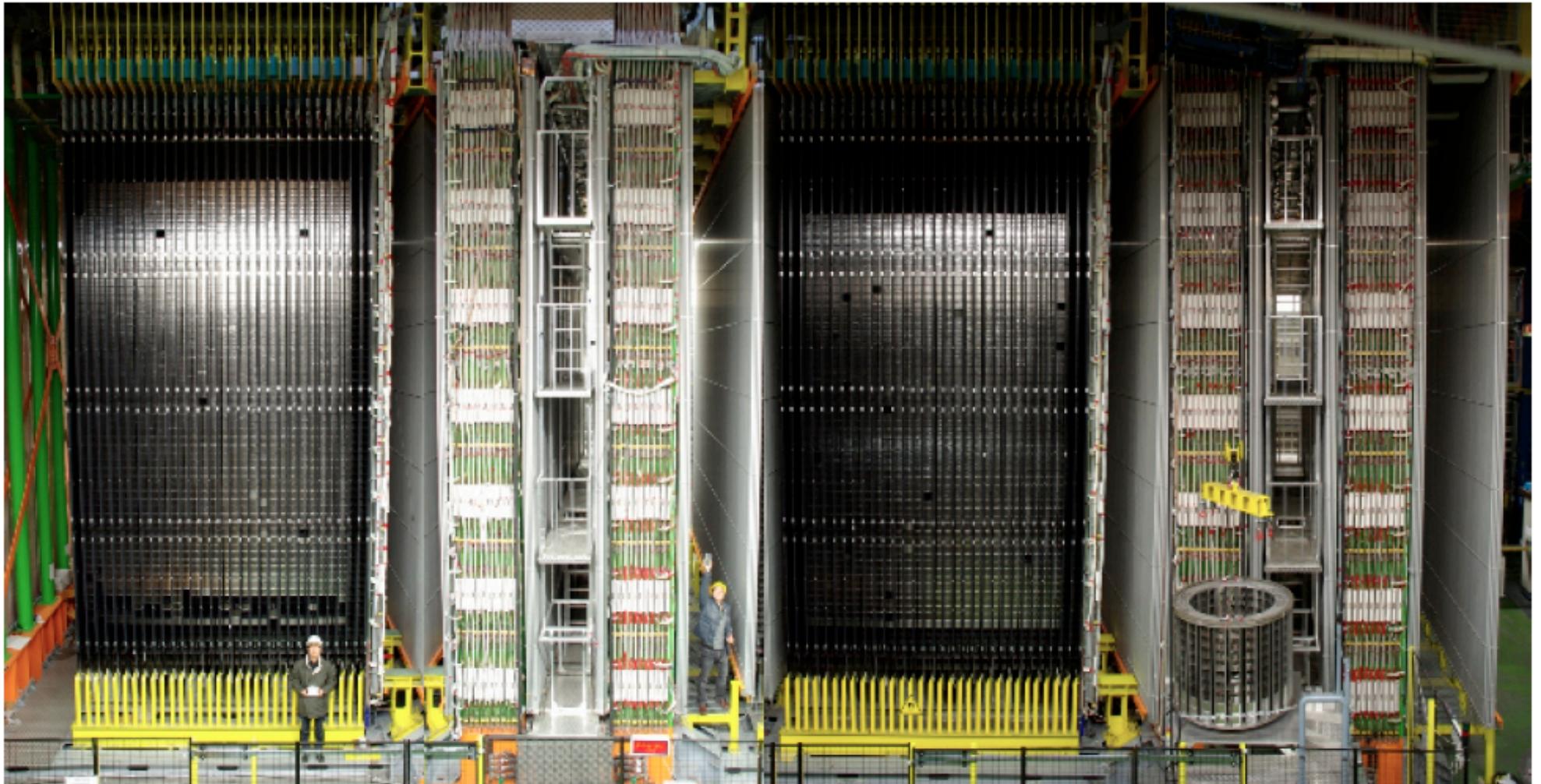
## Liquid Scintillator (Dbl) Chooz and KamLAND



Neutrino  
Detectors

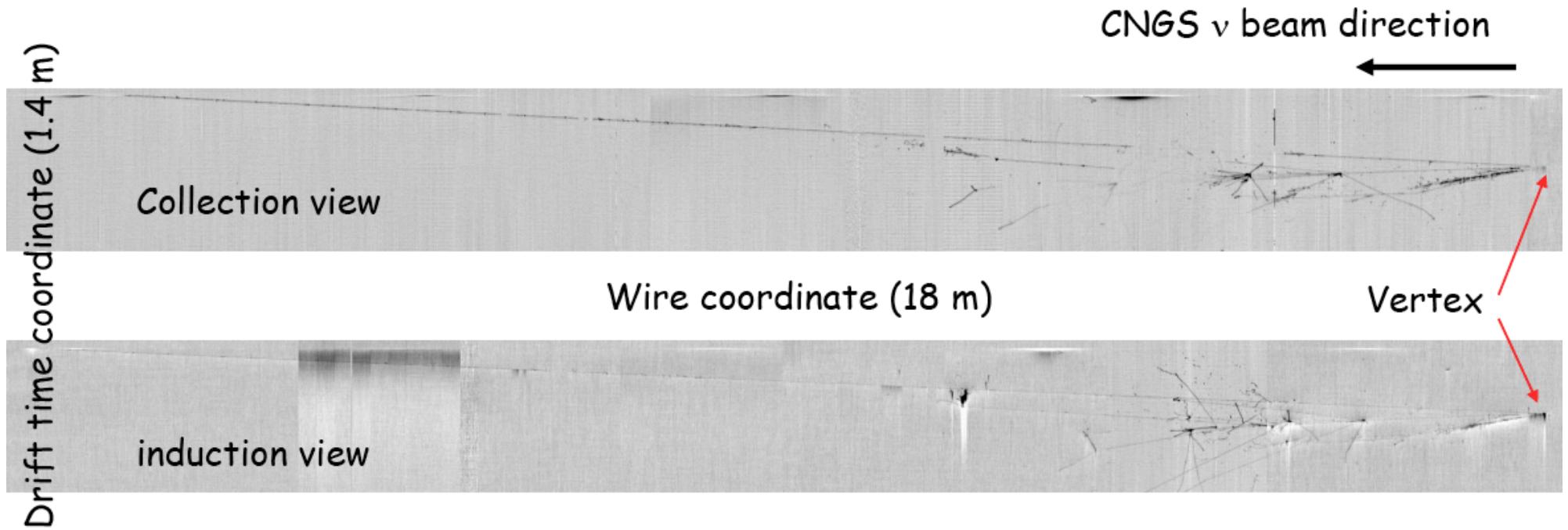


# Opera



- ❖ Fine grained “bricks” (1mm thick lead/emulsion).
- ❖ Electronic detectors point back at neutrino vertex.

# The first CNGS neutrino interaction in ICARUS T600



- Leading muon (crossing horizontally the whole cryostat)
- Two charged particle tracks undergoing hadronic interactions
- Two  $\gamma$  converting at 14 and 16 cm from vertex ( $\pi^0$ ?)
- Vertex not fully visible in collection view, due to locally wrong wire biasing

Neu2010

Slide# : 25